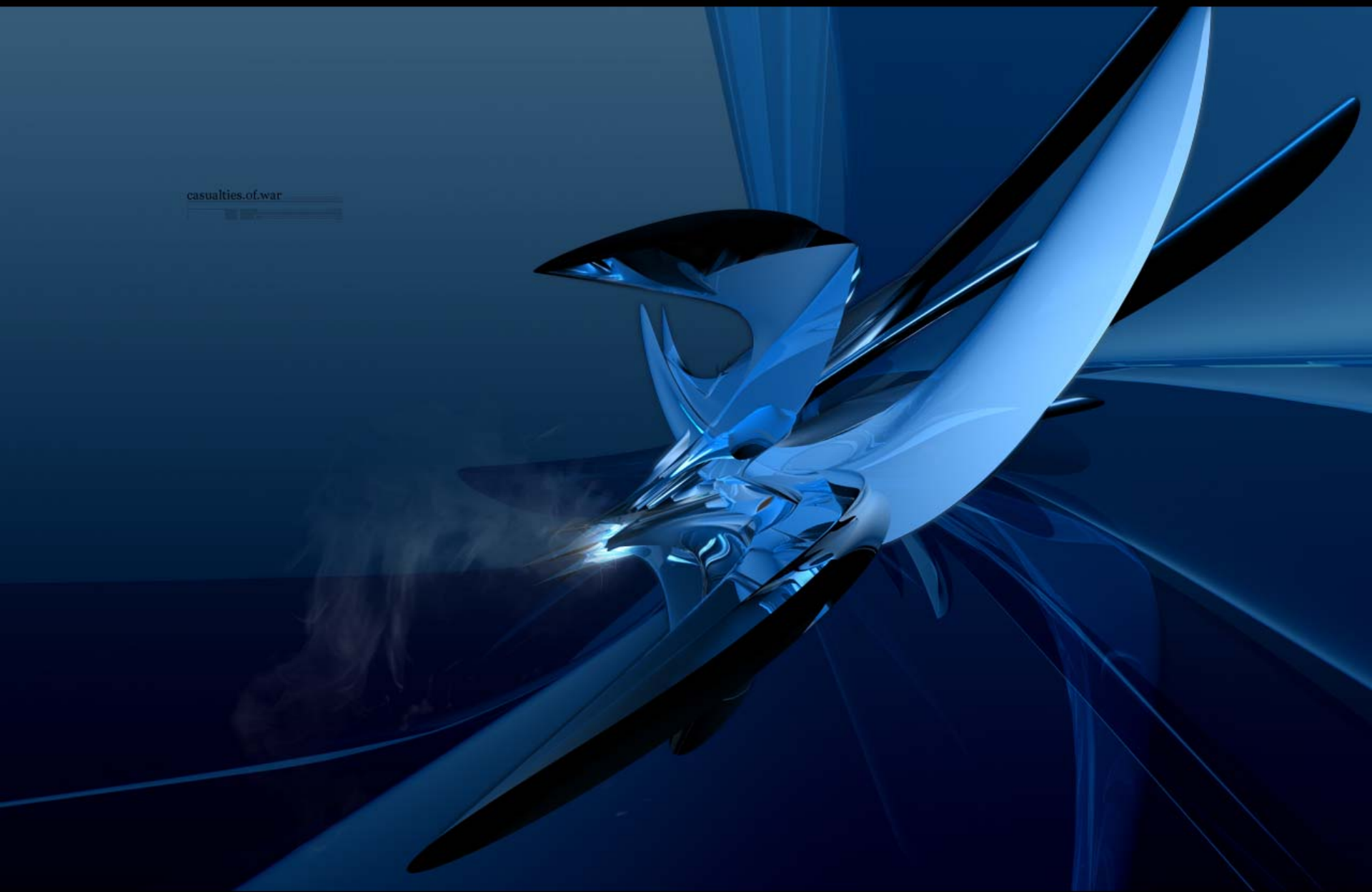


casualties.of.war



European Organization for Nuclear Research



CINVESTAV – Campus Mérida



Electron Cloud Effects in the LHC

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The background of the slide is a complex, abstract pattern of light trails. It features a dense network of thin, glowing lines in shades of blue, purple, and yellow, set against a dark, almost black background. The lines appear to be moving or vibrating, creating a sense of dynamic energy and depth. The overall effect is reminiscent of a microscopic view of a complex material or a digital data visualization.

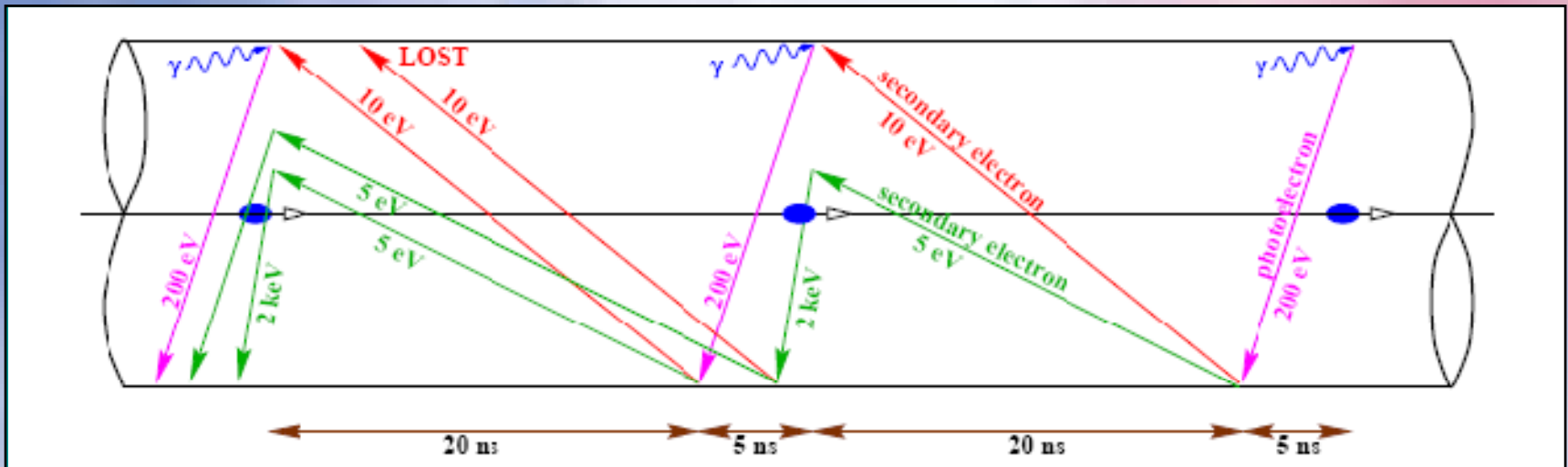
Outline

Introduction



Electron cloud build-up

The synchrotron radiation in the LHC creates a continuous flow of photoelectrons. These electrons are accelerated by the electric field of the bunch and hit the vacuum chamber where they create secondary electrons.



Photoemission, residual gas ionization and secondary emission give rise to a quasi-stationary electron cloud inside the beam pipe !!!

Electron cloud effects:

Due to e- induced gas desorption from the walls of the beam screen the vacuum pressure is increased by several orders of magnitude.

The electrons near the center of the vacuum chamber are attracted by the electric field of the beam and accumulate (“pinch”) inside the proton beam during a bunch passage. They can cause beam instabilities, emittance growth, even beam loss, and poor lifetime.

The energetic electrons heat the surfaces that they impact. Only a limited cooling capacity is available for the additional heat load due to the electron cloud.

Simulation code:

ECloud

enzyme

ECloud simulates the build up of the electron cloud.

- The ECLLOUD simulation includes the electric field of the beam, arbitrary magnetic fields, the electron space charge field, and image charges.
- As input numbers, the code requires various beam parameters, surface properties, the vacuum chamber geometry and the type of magnetic field



Therapy Sessions

Methodology

We made 3 sets of simulations:

Set A “nominal”	
Yield	Bunch spacing
1.1 - 1.7	25 ns
Nb	
2 x 10 ¹⁰ - 18 x 10 ¹⁰	

Set B “50-ns alternative”	
Yield	Bunch spacing
1.1 - 1.7	50 ns
Nb	
2 x 10 ¹⁰ - 18 x 10 ¹⁰	

Set C “upgrade”	
Yield	Bunch spacing
1.1 - 1.7	50 ns
Nb	
20 x 10 ¹⁰ - 60 x 10 ¹⁰	

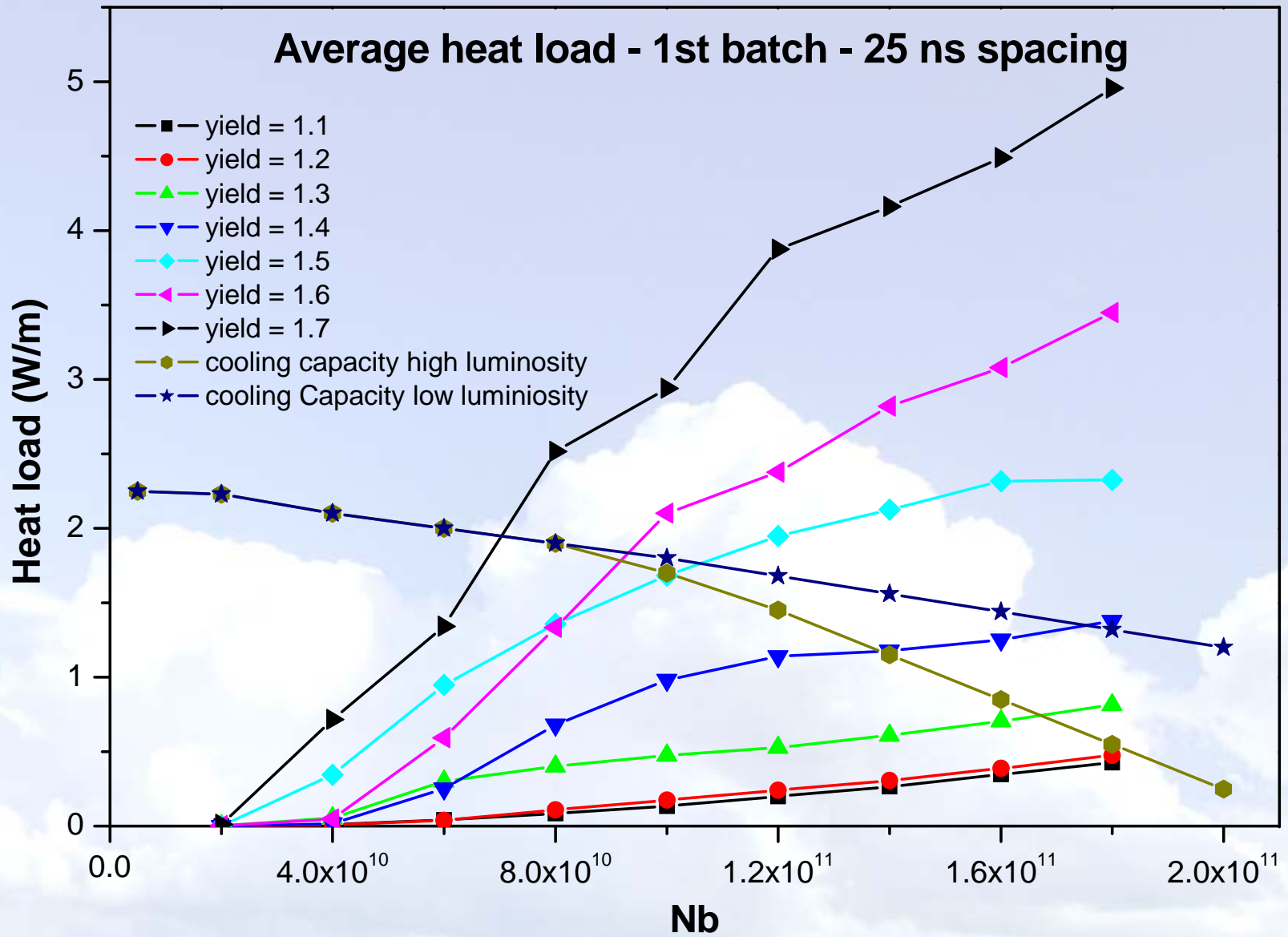
* For drift and dipole magnets the number of “bunch places” was 160 (2 trains) and for quadrupole we considered only 130 (50 bunches in second train).

IMAGINE YOU'LL FIND A NEW STATE OF MIND, ADDING LIGHT THROUGH YOUR CAREER, RIGHT?

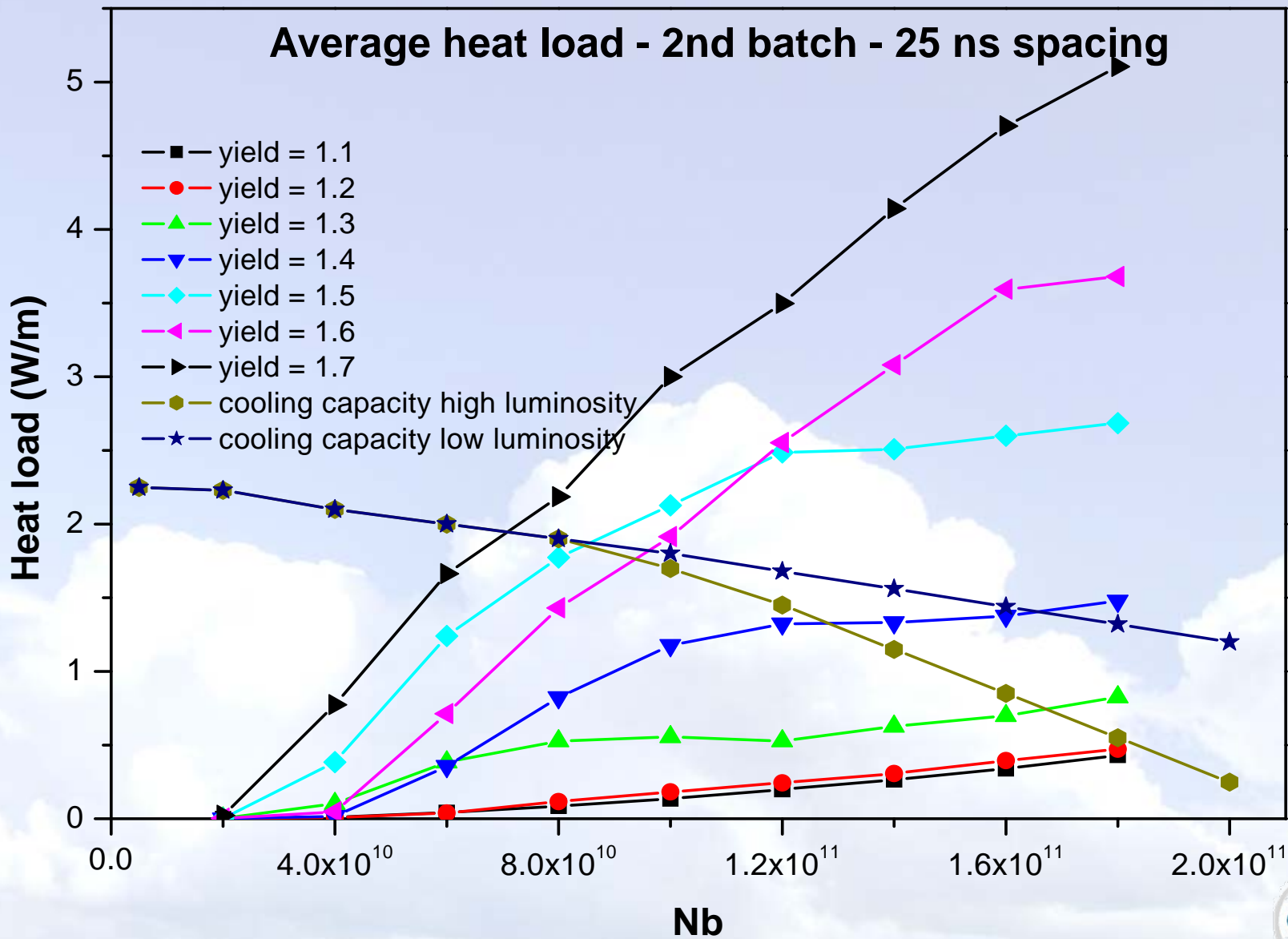
Results

IMAGINE YOU'LL FIND A NEW STATE OF MIND, ADDING LIGHT THROUGH YOUR CAREER, RIGHT?
THE FUTURE IS BEAUTIFUL

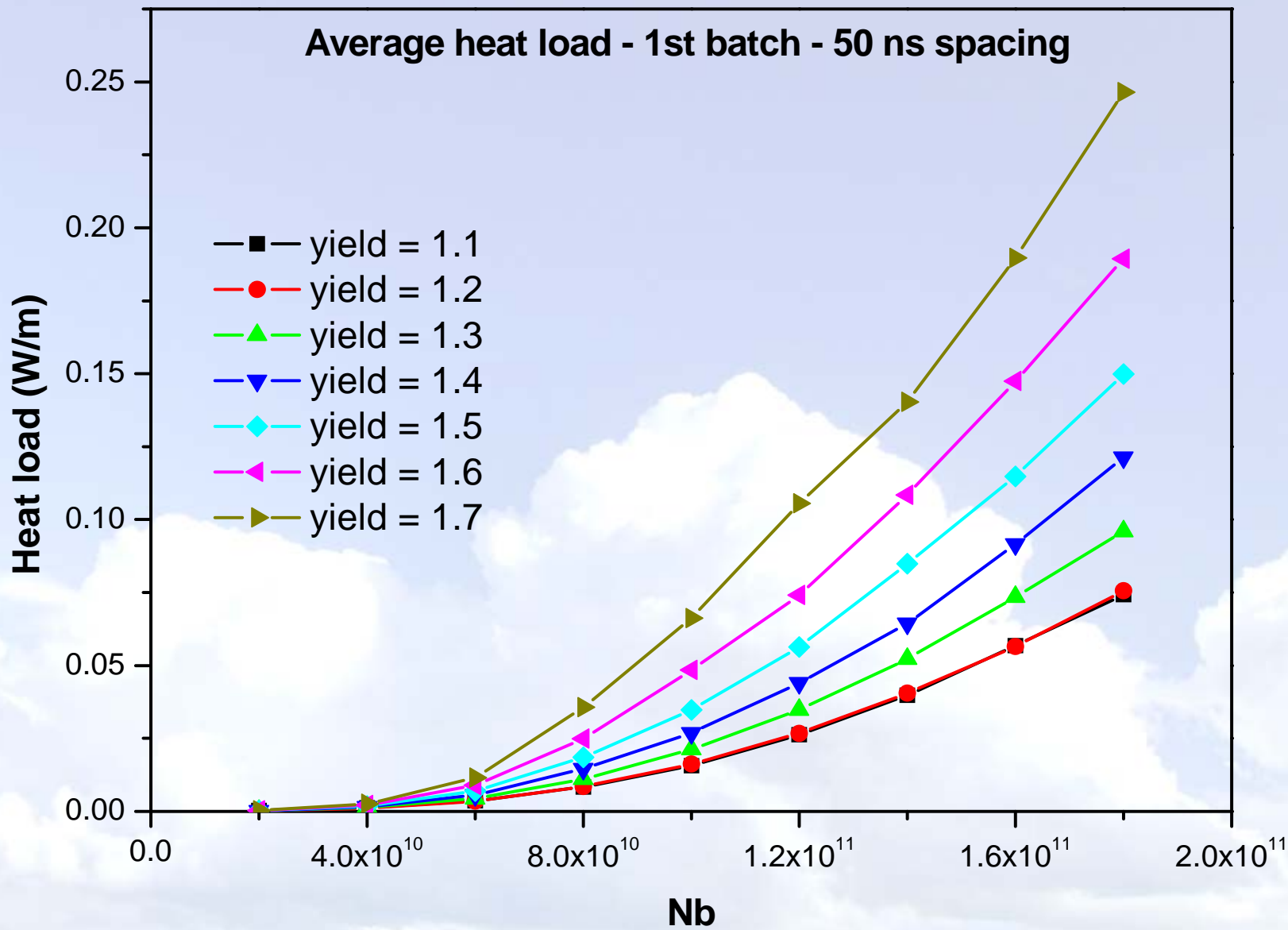
Average heat load - 1st batch - 25 ns spacing

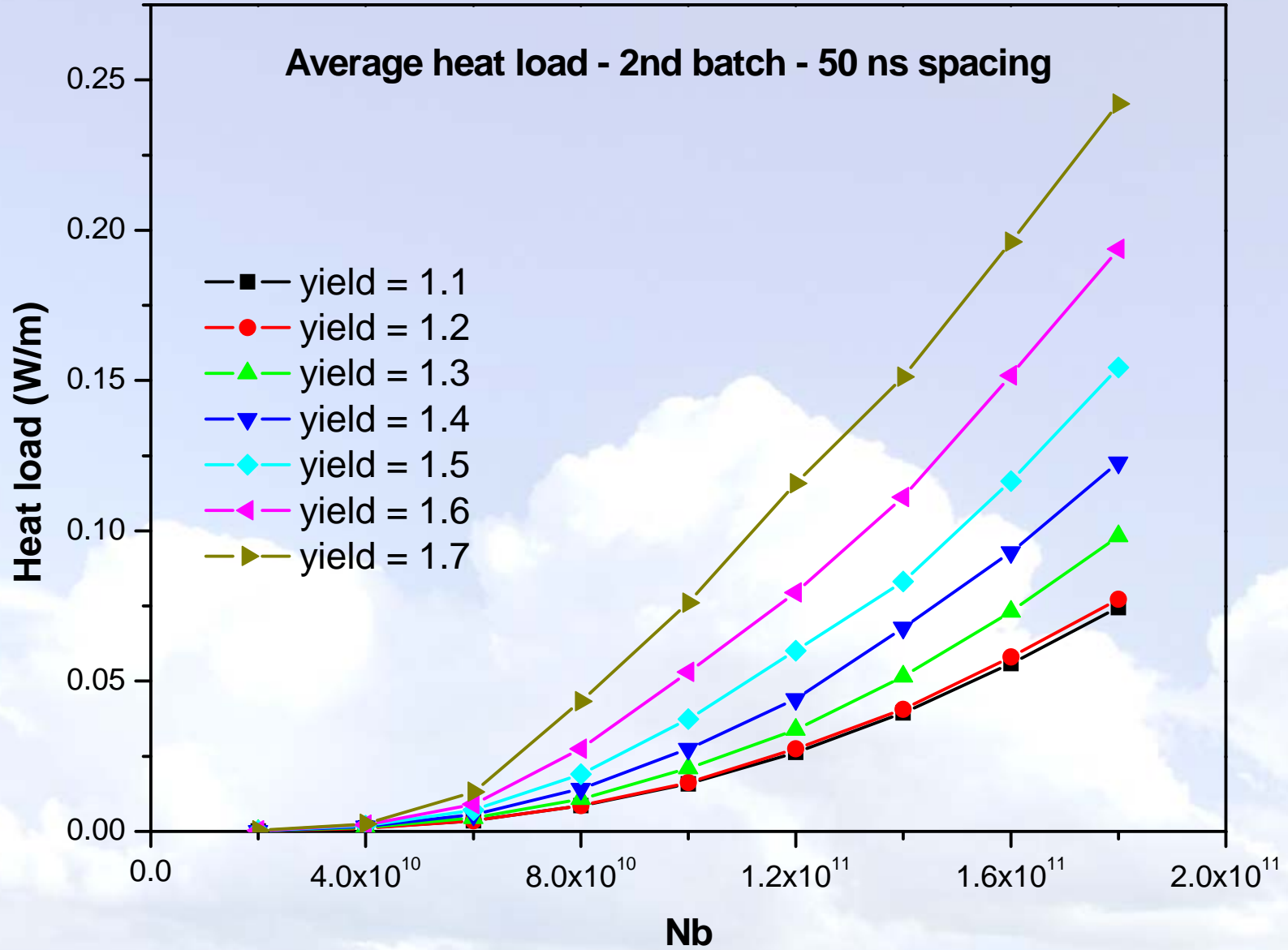


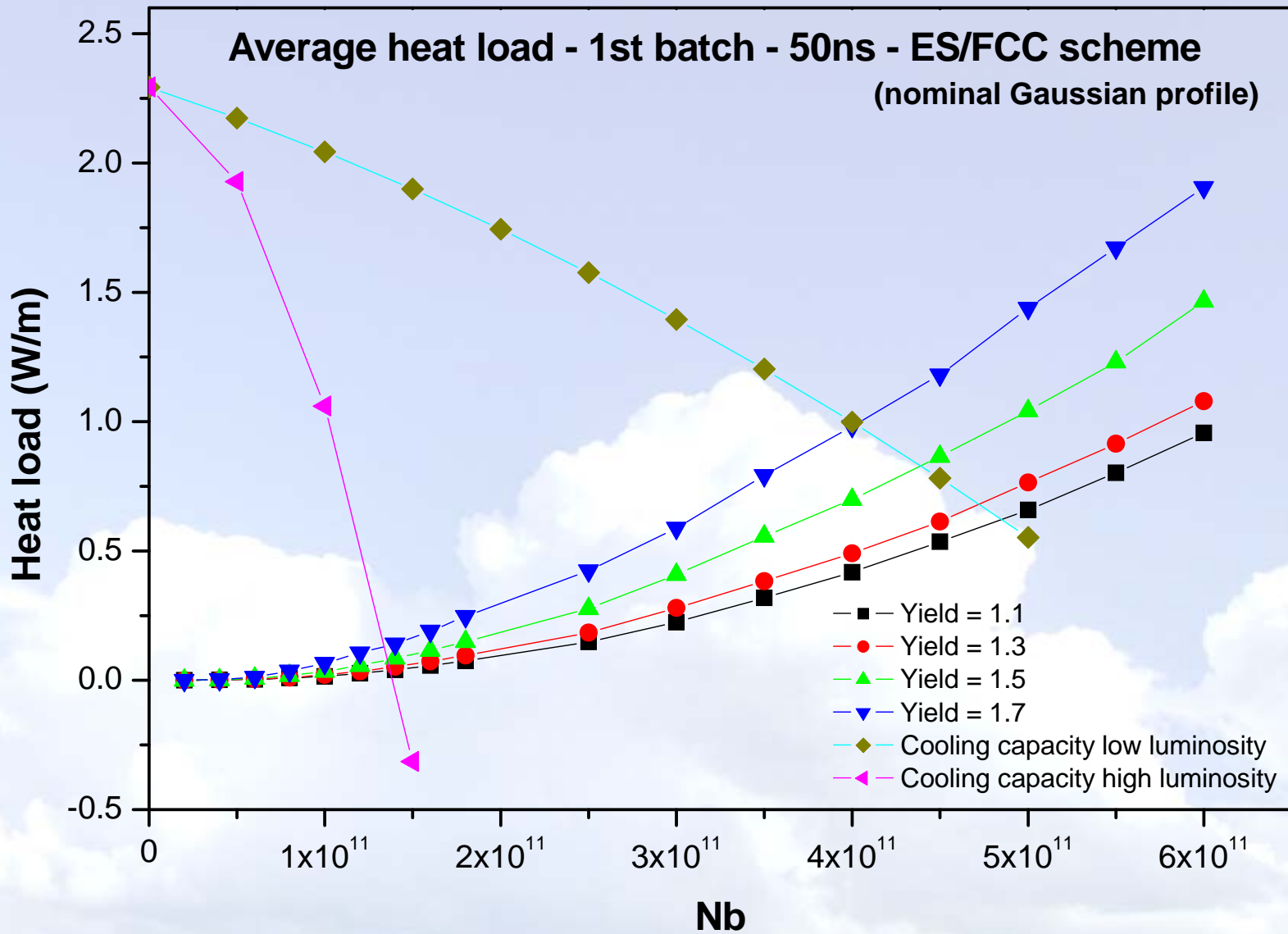
Average heat load - 2nd batch - 25 ns spacing

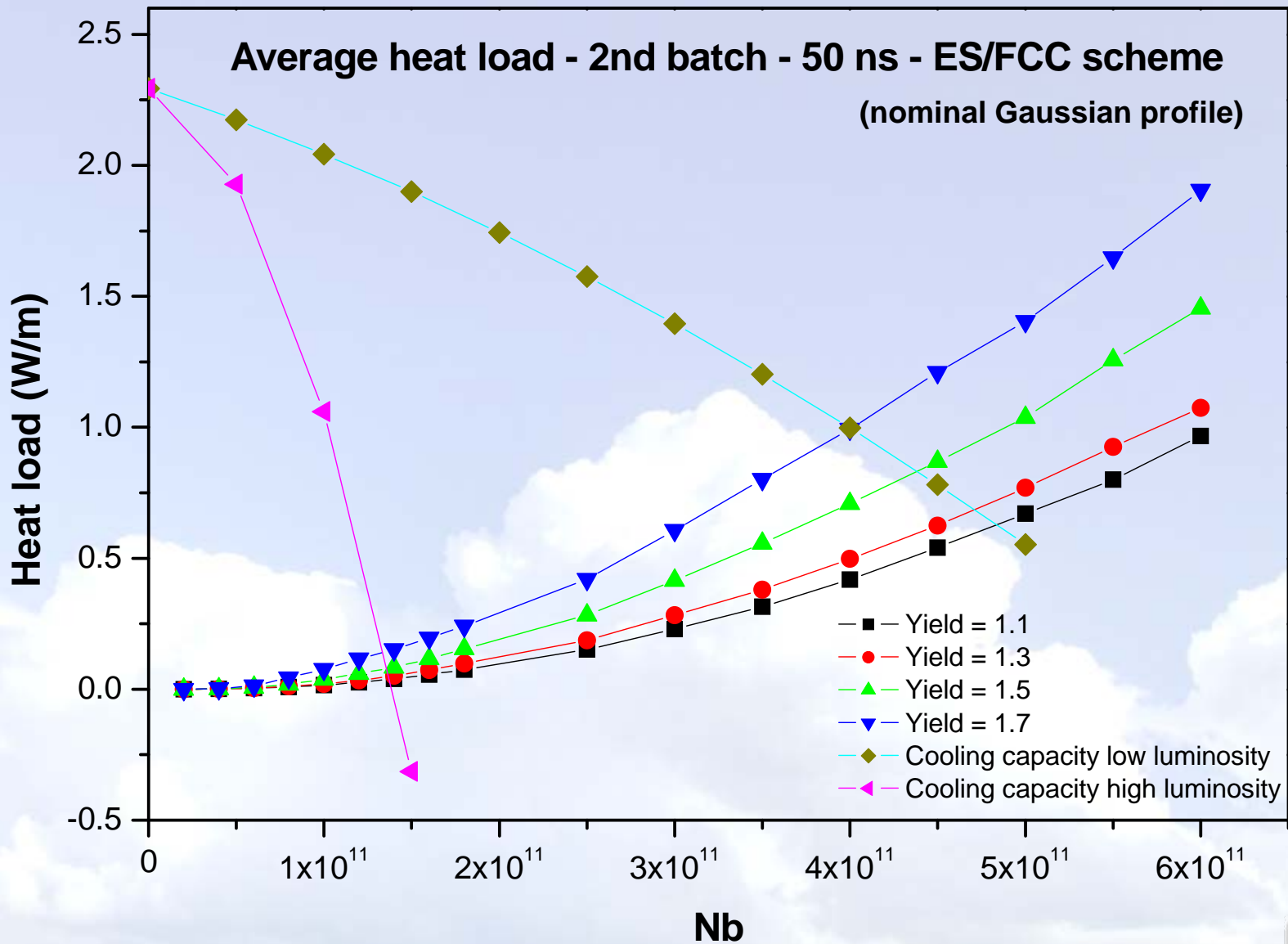


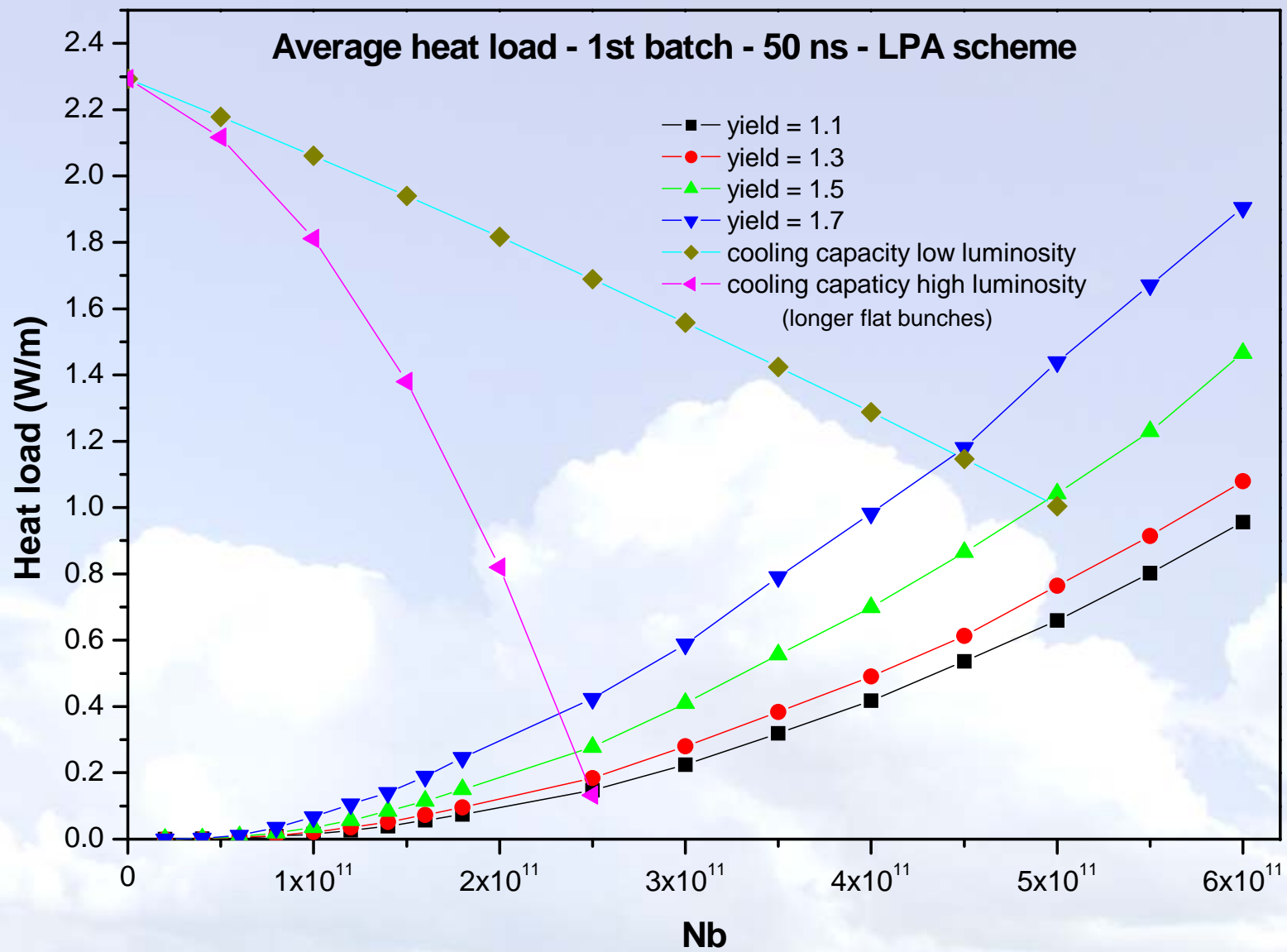
Average heat load - 1st batch - 50 ns spacing

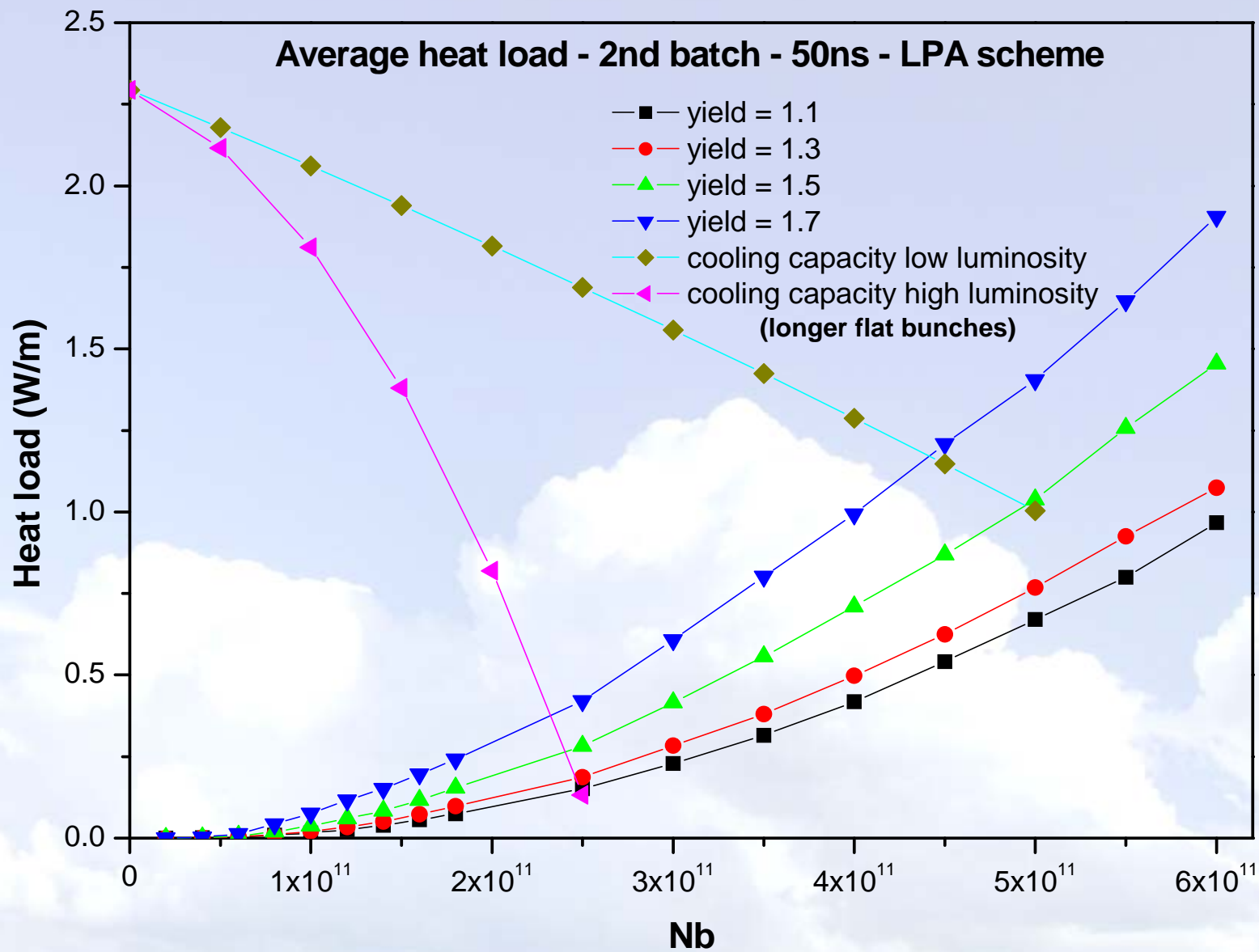














Conclusions

Conclusions:

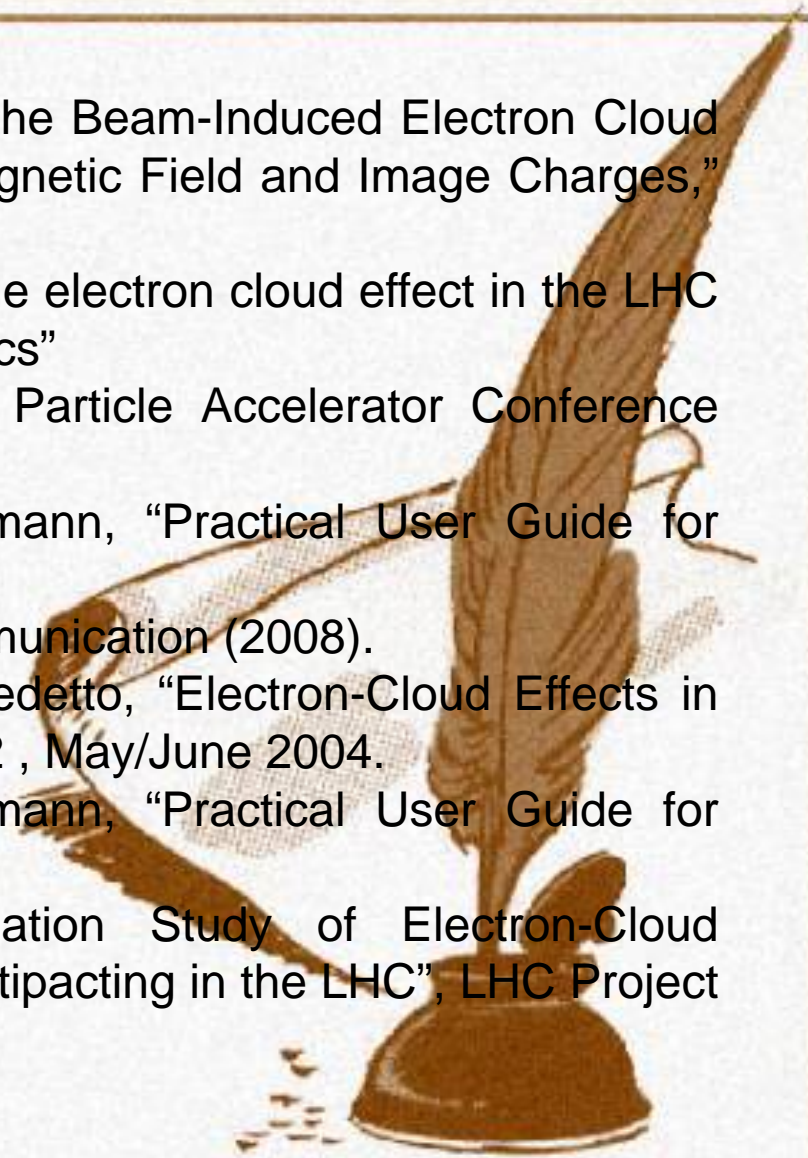
- heat load for 1st & 2nd batch almost the same
- 25 ns spacing:
 - for SEY < 1.3 ultimate parameters,
 - for SEY < 1.4 nominal LHC,
 - for SEY < 1.5 up to $N_b=9 \times 10^{10}$
- 50 ns spacing:
 - for nominal $\beta^*=0.55$ m up to $N_b > 2 \times 10^{11}$
- High-luminosity upgrade requires separate cooling for IRs; then
 - ES/FCC ($\beta^*=0.08$ m) up to $N_b \sim 4.5 \times 10^{11}$
 - LPA ($\beta^*=0.25$ m) up to $N_b \sim 5.5 \times 10^{11}$

Future work

- ❖ Compare heat load for Gaussian bunches with $\sigma_z=7.55$ cm and longer flat bunches with $l_b=41$ cm.
- ❖ Simulate PS and SPS experiments (later).
- ❖ Compare real LHC data with simulation (next year!?).

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*Thank you so much
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